# Emergent bubbling geometries in the plane wave matrix model 

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## Plane Wave Matrix Model (1D Quantum Mechanics)

## IIA SUGRA solution

This gravity solution was constructed by Lin and Maldacena in ' 05 .

Q: How does the space-time geometry emerge in the framework of the corresponding gauge theory?

## Plane Wave Matrix Model

 (1D Quantum Mechanics)
## IIA SUGRA solution

Geometry
Range of the eigenvalues $\longleftrightarrow$ Typical geometric scale

I'm going to show
how the geometry can be constructed from the Plane Wave Matrix Model!

## Plan of Talk

1) Gravity Side
2) PWMM (Gauge Theory)
3) Geometry from PWMM
4) Summary, Some Notes and Future Works

## Gravity Side

The IIA SUGRA solutions dual to $S U(2 \mid 4)$ symmetric theories have $R \times S O(3) \times S O(6)$ isometry.

The metric can be written by a single function $V(r, z)$ :
[Lin-Lunin-Maldacena '04, Lin-Maldacena '05]

$$
\begin{aligned}
& d s_{10}^{2}=\left(\frac{\ddot{V}-2 \dot{V}}{-V^{\prime \prime}}\right)^{1 / 2}\left\{-4 \frac{\ddot{V}}{\ddot{V}-2 \dot{V}} d t^{2}-2 \frac{V^{\prime \prime}}{\dot{V}}\left(d r^{2}+d z^{2}\right)+4 d \Omega_{5}^{2}+2 \frac{V^{\prime \prime} \dot{V}}{\Delta} d \Omega_{2}^{2}\right\}, \\
& C_{3}=-4 \frac{\dot{V}^{2} V^{\prime \prime}}{\Delta} d t \wedge d \Omega_{2}, \quad B_{2}=\left(\frac{\left(\dot{V}^{2}\right)^{\prime}}{\Delta}+2 z\right) d \Omega_{2}, \cdots \quad\left(\Delta=(\ddot{V}-2 \dot{V}) V^{\prime \prime}-\left(\dot{V}^{\prime}\right)^{2}\right)
\end{aligned}
$$

EoM and Killing spinor eq. for the geometry
$V$ satisfies the Laplace eq. in a 3D axisymmetric system.
Reducing to axisymmetric electrostatic problems with "conducting disks."

General geometry:

The geometries are
labeled by $\left\{N_{2}{ }^{(s)}, N_{5}{ }^{(s)}\right\}_{s=1, \ldots, \Lambda}$

$\Lambda=1$ case:
$V(r, z)=V_{0}\left(r^{2} z-\frac{2}{3} z^{3}\right)+\beta(\kappa) V_{0} R^{3} \int_{-R}^{R} d t\left(-\frac{1}{\sqrt{r^{2}+(z+d+i t)^{2}}}+\frac{1}{\sqrt{r^{2}+(z-d+i t)^{2}}}\right) \frac{f_{\kappa}(t)}{\pi}$.
$V_{0}$ : constant, $R$ : radius of disk, $\kappa=d / R, \beta(\kappa)$ : function of $\kappa$

$$
f_{\kappa}(x)-\frac{1}{\pi} \int_{-1}^{1} d y \frac{2 \kappa}{4 \kappa^{2}+(x-y)^{2}} f_{\kappa}(y)=1-\frac{2 \kappa}{\beta(\kappa)} x^{2}
$$



## PWONM (Gauge Theary Side)

PWMM is a 1D Matrix Theory with $S U(2 \mid 4)$ symmetry.
[Berenstein-Maldacena-Nastase '02]

$$
\begin{aligned}
& S=\frac{1}{g^{2}} \int d \tau \operatorname{Tr}\left(\frac{1}{2}\left(D_{\tau} X_{a}\right)^{2}+\frac{1}{2}\left(D_{\tau} X_{m}\right)^{2}+\frac{1}{4}\left(2 \epsilon_{a b c} X_{c}-i\left[X_{a}, X_{b}\right]\right)^{2}-\frac{1}{2}\left[X_{a}, X_{m}\right]^{2}\right. \\
& D_{\tau}=\partial_{\tau}-i\left[A_{\tau}, *\right]\left.-\frac{1}{4}\left[X_{m}, X_{n}\right]^{2}+\frac{1}{2} X_{m} X^{m}+\text { fermions }\right)
\end{aligned}
$$

$$
a, b=1,2,3, \quad m, n=4, \ldots, 9
$$

The vacua are given by $S U(2)$ rep.

$$
\hat{X}_{a}=-2 \bigoplus_{s=1}^{\Lambda}\left(\mathbf{1}_{N_{2}^{(s)}} \otimes L_{a}^{\left[N_{5}^{(s)}\right]}\right) \quad \begin{aligned}
& L_{a}^{\left[N_{5}(s)\right]: \operatorname{spin}\left(N_{5}^{(s)}-1\right) / 2 \text { rep. mat. }} \\
& N_{2}(s): \text { multiplicity of the spin }\left(N_{5}^{(s)}-1\right) / 2 \text { rep }
\end{aligned}
$$

-• labeled by $\left\{N_{2}{ }^{(s)}, N_{5}{ }^{(s)}\right\}_{s=1, \ldots, \Lambda}$

## So as to construct geametries

Consider simple vacua ( $\Lambda=1$ ) from now on.


Then apply the localization method to PWMM and get the same equation as that obtained on the gravity side:

$$
f_{\kappa}(x)-\frac{1}{\pi} \int_{-1}^{1} d y \frac{2 \kappa}{4 \kappa^{2}+(x-y)^{2}} f_{\kappa}(y)=1-\frac{2 \kappa}{\beta(\kappa)} x^{2}
$$

## Lacalization

Let us apply the localization method to PWMM. [rA.-shiki-OKada-Shimasaki '12]

- SUSY: quarter BPS sector such that

$$
\phi(\tau)=2\left(-X_{3}(\tau)+\sinh (\tau) X_{8}(\tau)+i \cosh (\tau) X_{9}(\tau)\right) \text { is invariant. }
$$

- B.C.: all fields approach to the vacuum configurations at $\tau \rightarrow \infty$.
- $\mathcal{V}: \delta_{s} \mathcal{V}$ is SUSY-invariant and positive-definite. $\quad \mathcal{V}=\int d \tau \operatorname{Tr}\left[\overline{\psi \delta_{s} \Psi}+\mathcal{V}_{\text {ghost }}\right]$

Then,

$$
Z(t):=\int \mathcal{D} X e^{-S[X]-t \delta_{s} \mathcal{V}} \quad \quad \frac{d Z(t)}{d t}=0 \quad \int \mathcal{D} X e^{-S[X]}=Z(0)=Z(\infty)
$$

Localized around $\delta_{s} \mathcal{V}=0$ !

$$
\hat{X}_{a}(\tau)=-2 \bigoplus_{s=1}^{\Lambda}\left(\mathbf{1}_{N_{2}^{(s)}} \otimes L_{a}^{\left[N_{5}^{(s)}\right]}\right) \quad \hat{X}_{9}(\tau)=\frac{M}{\cosh (\tau)} \quad\left(\left[L_{a}, M\right]=0\right)
$$

## Lacalization

Hence, expectation values of any supersymmetric operators can be obtained exactly by 1 -loop calculations.

$$
\left\langle\prod_{a} \operatorname{Tr} f_{a}\left(\phi\left(\tau_{a}\right)\right)\right\rangle=\left\langle\prod_{a} \operatorname{Tr} f_{a}\left(4 L_{3}+2 i M\right)\right\rangle_{M M}
$$

When $\Lambda=1,\langle \rangle_{M M}$ is evaluated by the following partition function:

$$
Z=\int \prod_{i} d q_{i} \prod_{J=0}^{N_{s}-1} \prod_{i>j}^{N_{2}} \frac{\left\{(2 J+2)^{2}+\left(q_{i}-q_{j}\right)^{2}\right\}\left\{(2 J)^{2}+\left(q_{i}-q_{j}\right)^{2}\right\}}{\left\{(2 J+1)^{2}+\left(q_{i}-q_{j}\right)^{2}\right\}^{2}} e^{-\frac{2 N_{s}}{g^{2}} \sum_{i} q_{i}{ }^{2}} .
$$

qi: eigenvalues of $Q$, where $M=Q \otimes \mathbf{1}_{N_{5}}$

## Interacting Fermi Gas System

In the large $-N_{5}$ limit, the partition fn. can be written as

$$
\begin{gathered}
Z=\int \prod_{i} d q_{i} \prod_{i>j}^{N_{2}} \tanh ^{2} \frac{\pi\left(q_{i}-q_{j}\right)}{2} \exp \left\{-\frac{2 N_{5}}{g^{2}} \sum_{i} q_{i}^{2}+\frac{2 N_{5}}{\left(2 N_{5}\right)^{2}+\left(q_{i}-q_{j}\right)^{2}}-\cdots\right\} \\
\text { Repulsive force } \quad \text { Central force } \quad \text { Attractive force }
\end{gathered}
$$



$$
\begin{aligned}
& \text { By using the Cauchy id., } \\
& \prod_{i \neq j}^{N_{2}} \tanh \frac{\pi\left(q_{i}-q_{j}\right)}{2}=\sum_{\sigma \in S_{N_{2}}}\left(-\epsilon^{\epsilon(\sigma)} \prod_{i=1}^{N_{2}} \frac{1}{\cosh \frac{\pi\left(q_{i}-q_{\sigma(i)}\right)}{2}}\right.
\end{aligned}
$$

[Marino-Putrov '12]
Interacting Fermi gas system with hamiltonian

$$
H=\sum_{i=1}^{N_{2}} \log \cosh p_{i}+\frac{2 N_{5}}{g^{2}} \sum_{i=1}^{N_{2}} q_{i}^{2}-\frac{1}{2} \sum_{i \neq j}^{N_{2}} \frac{2 N_{5}}{\left(2 N_{5}\right)^{2}+\left(q_{i}-q_{j}\right)^{2}}
$$

## Interacting Fermi Gas System

We consider the limit in which the SUGRA approx. is valid: large- $N_{2}$, large- $N_{5}, \lambda:=g^{2} N_{2} \gg N_{5}$

- . Thomas-Fermi approximation at zero temperature

The system is described by mean-field density $\rho(q)$.

$$
=\text { eigenvalue density }
$$

It is determined by the following saddle point eq.:

$$
\mu=\pi \rho(q)+\frac{2 N_{5}}{g^{2}} q^{2}-\int_{-q_{m}}^{q_{m}} d q^{\prime} \frac{2 N_{5}}{\left(2 N_{5}\right)^{2}+\left(q-q^{\prime}\right)^{2}} \rho\left(q^{\prime}\right)
$$

$\mu$ : chemical potential
$q_{m}$ : upper limit of the support of $\rho(q)$

Construction of geametry

$$
\frac{\pi}{\mu} \rho\left(q_{m} x\right)-\int_{-1}^{1} d y \frac{2 N_{5} / q_{m}}{\left(2 N_{5} / q_{m}\right)^{2}+(x-y)^{2}} \frac{\rho\left(q_{m} y\right)}{\mu}=1-\frac{2 N_{5} q_{m}^{2}}{\mu g^{2}} x^{2}
$$

$$
f_{\kappa}(x)-\frac{1}{\pi} \int_{-1}^{1} d y \frac{2 \kappa}{4 \kappa^{2}+(x-y)^{2}} f_{\kappa}(y)=1-\frac{2 \kappa}{\beta(\kappa)} x^{2}
$$

Construction of geametry

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$$

Identification to obtain the gravity picture

$$
\rho(q)=\frac{\mu}{\pi} f_{\kappa}\left(q / q_{m}\right) \quad \propto \text { charge density on the disk }
$$

By this identification, the range of the eigenvalue dist. can be written by the quantities on the gravity side:

$$
q_{m}=\frac{2}{\pi} R \quad \text { so } \quad R_{S^{5}}^{2}=2 \pi q_{m}
$$

The gravity solution corresponding to a simple vacuum in PWMM has been constructed from the gauge theory side!

## D2-and NS5-brane limit

On the gravity side, these two double scaling limits are known:

D2-brane limit


$$
R_{S^{5}}^{2}=\pi\left(3 g_{S^{2}}^{2} N_{2}\right)^{\frac{1}{3}}
$$

NS5-brane limit

$R_{S^{5}}^{2}=2 \pi(8 \lambda)^{\frac{1}{4}}$

On the gauge theory side, solutions in the corresponding limits are

$$
\rho(q)=\frac{\mu}{\pi}\left[1-\left(\frac{q}{q_{m}}\right)^{2}\right], \quad q_{m}=\left(\frac{3 \pi \lambda}{8 N_{5}}\right)^{\frac{1}{3}} \quad \rho(q)=\frac{\mu q_{m}}{3 \pi N_{5}}\left[1-\left(\frac{q}{q_{m}}\right)^{2}\right]^{\frac{3}{2}}, q_{m}=(8 \lambda)^{\frac{1}{4}}
$$

Then, $R_{S^{5}}^{2}=2 \pi q_{m}$
reproduce the same results as those on the gravity side.

## Summary

- The mean-field density, that is, the eigenvalue distribution in the matrix integral determines the geometry.
- The range of the eigenvalue dist. corresponds to $S_{5}$ radius.


## Same Notes and Future Works

- We can also reconstruct geometries dual to general vacua ( $\Lambda \neq 1$ ) in the same way.
- This method also works for other $S U(2 \mid 4)$ symmetric theories such as SYM on $R \times S^{2}$ and $R \times S^{3} / Z_{k}$.
- Exact proof for the NS5-brane limit. (Work in progress)

